

**Annual Report – 1/12/2005 – 30/11/2006<sup>1</sup>**

**for**

**Research Infrastructure RTD Project**

**ASTRO-WISE  
HPRI-CT-2001-50029**

**Groningen, 1 Februari 2007**

**NOVA/ OmegaCEN  
Kapteyn Instituut  
Rijksuniversiteit Groningen**

This document can be found on the INTERNET at the following address:

<http://www.astro-wise.org/Public/Aweannual-2006.pdf>

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***The Annual Report covers the fifth year of the project.***



## 1. EXECUTIVE SUMMARY

This summary should be as comprehensive and clear as possible. It should contain, in about 200 words a description:

- of the main objectives and characteristics of the RTD project, emphasising its innovative aspects and its potential impact on the quality and quantity of access provided by Europe's research infrastructures, and
- of the work performed and achievements.

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### Objectives

The ASTRO-WISE programme co-ordinates the development work to deal with and access astronomical wide field imaging data. It provides an operational environment for researchers to analyse and access the huge data volumes of observational data produced by a new generation of Europe's wide field imaging telescopes.

The programme pools the expertise of a number of European groups with experience or commitment to optical wide-field survey work. All partner sites will function as (linked) national data centres. The code and expertise to run additional satellite centres is a deliverable of the project.

ASTRO-WISE develops and disseminates in the community software tools needed to access the wide-field image data, to perform individual research programmes. The huge data volumes require an innovative dynamical approach, in which results can be re-derived, customized to the users specific needs. To this end, the various calibration data and other input files are distributed over a network, which connects the data centres. The geographical distribution of the key information (both methods and data) is provided by the system, allowing the various National Data centres to work cooperatively in an efficient manner and optimally profit from the work done at each of the centres. This enhancement, consolidation and dissemination through National data centres, is going to build up the Astronomical Wide-field Imaging System for Europe.

### Results

In this fifth year the project plan as outlined in the contract has been successfully completed. The entire information system has been built and is distributed over the community. The system has been qualified using astronomical wide field imaging data from the [WFI@2.2m](#) telescope, the Wide field imager at La Palma (Canary Islands), test data of OmegaCAM, the MDM telescope (US) and the Suprime-Cam Camera at the 8m Subaru telescope (Hawaii). Extensive web services have been put into operations. The great success of the e-science aspects of the system triggered other parties to use it as a platform: this includes the Coma Legacy Survey with the ACS camera at the Hubble Space telescope, radio imaging data from the Westerbork radio telescope and even scans of the Dutch National Archive of the handwritten archives of the Dutch Government which are processed with Artificial Intelligence techniques to digitally archive this Cultural Heritage.

National AstroWise data center nodes were put into operation at Groningen, Munich, Napoli and Paris, while AstroWise satellite nodes have been put into place at Leiden and Bonn, also serving research groups in Nijmegen, Bochum, Heidelberg and Santiago (Chile). The system i.e. code base, documentation and data is ready for delivery to other parties on request.

The AstroWise information system has been presented at several important conferences like the IAU General Assembly in Praha, the ADASS XVI in Tucson (USA). Papers describing the system have been published (Valentijn et al. 2006, Valentijn & Verdoes Kleijn 2006). Particular the novel concepts of "entire backward chaining", i.e the linking of data products to the input raw data draw attention and provides a unique backbone for the infrastructure for advanced distributed e-science.

The system is ready for deriving, hosting, and warehousing the data of new generations of massive data accumulators and is scalable up to the Petabyte data regime.

Advanced courses were held to introduce the user community to the system: a centralized course at the Lorentz center in Leiden (40 persons) and local courses at the partner sites München, Bonn and Napoli, typically attended by 15-25 persons.

Papers describing and/or using the Astro-WISE Information System (see <http://www.astro-wise.org/publications.shtml> for full articles):

- Valentijn & Verdoes Kleijn 2006, "The Astro-WISE system: a federated information accumulator for astronomy", ERCIM News, 65, 20
- Valentijn et al., 2006, "Astro-WISE: Chaining to the Universe", ADASS XVI, ASP Conference Series, 2006, Shaw, Hill, Bell, eds, in press
- Sikkema et al, 2006, "Globular cluster systems of six shell galaxies", Astronomy & Astrophysics, 458, 53
- Valentijn & Begeman, 2006, "Creating and analysing multi-dimensional data in Astro-WISE environment: a test case", white paper
- Verdoes Kleijn et al., 2006, "The secondary standards programme for OmegaCAM at the VST", Standardization workshop, ASP Conference series, Sterken editor, in press
- Kovac, 2006, "Searching for the lowest mass galaxies: an HI perspective", PhD thesis
- Pavlov et al, 2006, "The VST data processing within the GRID", Computational GRIDs for astrophysics in Italy, in press

On 14 November 2005 AstroWise has publicly announced its web portal [www.astro-wise.org/portal](http://www.astro-wise.org/portal) which contains various web services, including code base viewers, database viewers, a calibration file monitor and updater, collected documentation, a menu with "HOW TOs" for users and an test version of a very innovative distributed processor facility, which includes a viewer to all data items dependencies and their actual status (OK or outdated) in the system. During the fifth year these services were significantly optimized and qualified according to the needs of the users..

Hardware for WP5 (disk storage) and WP4 (computers) have been upgraded at various nodes.

**Project Manager** contact information- Valentijn see previous page

## 2. COMPARISON WITH THE PROJECT PROGRAMME

### Scientific and Technical Performance

This section should be an overview covering all the aspects of the work undertaken by the partners to date.

#### 2.1 Summary of the specific project objectives

The programme will establish a number of data centres where the voluminous wide-field imaging data can be processed to the specifications of the user, who can select on the basis of atmospheric conditions, quality of calibration data, etc which data to process and how. The linking of the data centres means that the user will always have access to the most up-to-date data.

highlights, if any, of important research results from the project (please, attach abstract of corresponding publication to Annex 1);

#### 2.2.1 Highlights

The official starting date of the project is 1/12/2001; so this report describes the fifth year of the project.

The programme executed the last phase: "Phase III: qualify, operate and populate the mass production system", highlights include:

- The full AstroWise e-science information system, its code base, webservices and documentation has been installed and delivered.
- The information system is connected to file servers and compute servers in a European wide grid, together forming an extensive e-science infrastructure.
- A public portal bundling an extensive set of GRID and web services has been operated and improved. It was opened on 14 November 2005 ([www.astro-wise.org/portal](http://www.astro-wise.org/portal)). Some services can be reached by the "world", while others are restricted to registered users.
- A tutorial/training week "Astro-Wise workshop, Preparing for Surveys" was organized by ASTRO-WISE and was attended by 40 persons of the consortium (Leiden Lorentz center 20-23 November 2006); additional training courses involving scientist were held in Napoli and Garching bei München.
- The system was further populated and qualified on a very large dataset of 10 Terabyte of wide field imaging data which included the Suprime-Cam camera at the Subaru telescope and WFI@2.2m telescope data (which was ingested during year 4) which:
  - was partly processed on the parallel clusters (provided by WP4)
  - by pipeline software delivered by WP1,
  - was viewed by users by AstroWise provided user interfaces (WP2) connected to the database (WP3)
  - was archived using the databases (WP3) connected to file servers and storage (WP5).
- The prototype image pipeline (WP1), written in the Astro-Wise paradigm, was extensively tested and improved using these test data and also that of other wide field imaging instruments (La Palma WFC, MDM).

- Major upgrade of web services- including extensive on-line documentation such as “How-Tos”- was completed
- A European wide federation of databases has been implemented
- Surface photometry packages have been implemented (GALPHOT, GALFIIT)
- Photometric Redshift Tool (PhotRed) and Light Variability Tools (VODIA, MDia) were improved and fully integrated into the system
- A image cutout server for fast web viewing and input to surface photometry was built and integrated in various services.
- The system is connected to the EuroVO Aladin viewer.
- Additional data sources of the radio telescope WSRT were ingested
- A pioneer project involving Artificial Intelligence, for pattern recognition of digital scans of handwritten text using super computers has been integrated and tested in the system.

The organization of the project, as outlined in the contract, was maintained:

- A PI meeting was held, on 22 Nov 2006 in Leiden (minutes are listed on [www.astro-wise.org/minutes/minutes-nov06.shtml](http://www.astro-wise.org/minutes/minutes-nov06.shtml)).
- 2 large meetings/workshops were held to review with developers and the user community together the functionalities of the system, leading to elaborate context domains (i.e projects and research groups, equivalent to Virtual Organizations) in the system
- The project web site [www.astro-wise.org](http://www.astro-wise.org) was maintained listing descriptions of the WPs, documentation pages, tutorials, agenda, minutes etc, etc. The project website is now upgraded for general use by the astronomical community.
- A project news group [news@astro-wise.org](mailto:news@astro-wise.org) and issues mailinglist (issues@astro-wise.org) is frequently used by all partners.
- teleconferences between local contacts were held monthly, and in addition weekly internet meetings between Data base administrators were continued.

The project organizational scheme with short communication lines both in between developers and with user/researchers was extremely useful during the qualifications and for achieving the user friendly-ness of the system. The other AstroWise typical organizational aspect, the central role of databases (WP3) in the whole system and connecting all the work packages has proven to be extremely useful.

Work packages (1-5) describe the project:

- WP1: design, data model and development rules for the data reduction pipeline to be built.
- WP2: Description of visualization tools and particularly in the case of the new to be developed visualization tool *detailed* design specifications are included. Development of database access tools.
- WP3: Design of Federated Database to distribute administrative and calibration data as well as documentation and software over the participating sites in support of WP1 and WP2. In addition the database will deliver or point to data items such as raw and processed image data and will contain source lists. WP3 provides the glue for the whole project.

- WP4-5 Lists the hardware currently operated or expected in the future by the partners. Pooling expertise for the setting up of the edge of technology hardware configurations (parallel processors – WP4 and dozens of Terabyte disk storage - WP5).

The Architectural Design Document (ADD) Version 1.0 forms the basis for the actual implementations, and constitutes the blueprint of the project. The implementation of the system was successfully finalized in the fifth year according the specifications in the ADD. Special attention was given to building web and GRID services for the system enhancing its e-science aspects and qualifying the system with small test projects.

Further highlights during the reporting period include:

- Meetings of complementary programmes were attended: SISCO Durham “end of SISCO conference” Aug 2006, “Euro-VO DCA board” 2-3 Sept Strasbourg, for coordination (WP6)
- The AstroWise information system has been presented at several important conferences like the IAU General Assembly in Praha, the ADASS XVI in Tucson (USA). Papers describing the system have been delivered( Valentijn et al 2006, Valentijn & Verdoes Kleijn 2006). In particular the novel concepts of “entire backward chaining”, i.e the linking of data products to the input raw data draw attention and provide a unique backbone for the infrastructure for advanced distributed e-science.

## 2.2.2 Technical Progress on a task by task basis

The main progress for each of the tasks WP1-5 is the implementation of the design as worded in the ADD. Here, we list activities on the implementation and qualification of the Astro-Wise Information system and GRID (hereafter Awe).

- **WP1 NOVA:** The prototype **image pipeline** has been further qualified with large amounts of data and was delivered. The Awe pipeline converts raw image data from the telescope into astrometrically and photometrically calibrated images. The database keeps track of all operations done to the data and eventually facilitates on-the-fly- re-processing of the data, user customized. NOVA/OmegaCEN extensively tested and upgraded the prototype pipelines in an operational environment using actual data from the WFI@2.2m instrument as test data. The new “target processing” facility (November 2005) was extensively tested, upgraded and delivered.
- **WP1 MPE/USM:** In order to test the Awe software from the first to the last step, and to familiarize a larger group of people with the AstroWise concepts as well as its operation from the user's point of view, a tutorial was organized and held from August until November (see [www.astro-wise.org/doc\\_usm.shtml](http://www.astro-wise.org/doc_usm.shtml) for tutorial material). The data used were observations of the GOODS field, taken with the WFI camera at La Silla in October of 2004, together with a set of corresponding calibration data and standard star observations. The organizing core group set up a database and ingested the WFI data. Accounts were created for access to the database. The tutorial consisted of a set of talks and tutored exercises. Although the web based target processing interface was also introduced, emphasis was laid in the tutorial on the use of the AstroWise Environment command prompt that gives the user access to the full set of features delivered by AstroWise. A smaller group then continued with the creation of SourceLists (Catalogs) and running the photometric redshift code. For general discussion and the documentation of problems that were encountered while gaining experience with the AstroWise system a forum software was installed on the MPE web server. The forum was opened up for registration to the groups in Groningen and Leiden after the AstroWise workshop in November 2006. The encountered problems were also posted to the AstroWise issues mailing list.
- **WP1 OAC:** Fine tuning and maintenance of the Astro-Wise pipeline on the OAC 17 nodes Opteron based beowulf system. Awe is fully operational at OAC. Testing the Awe pipeline: new tests were done using the OACDF data (see Alcalà et al. 2004, A&A 428, 339; <http://www.na.astro.it/oacdf-bin/cdfcgi> for more details), in particular to check astrometric and photometric performance.
  - Astrometry: global astrometry was obtained with final accuracy of the order of 0.05 arcsec.
  - Photometry: new tests on gain correction and PSF quality in the final coadded image using the "flux\_radius" parameter of SExtractor to separate the star sequence from the background galaxies.
- **WP1 (OAC)** A tutorial was organised at OAC in order to present the Awe environment to the users with some practical examples of data reduction. All presentations and exercises are available at the AstroWise web site ([http://www.astro-wise.org/doc\\_oac.shtml](http://www.astro-wise.org/doc_oac.shtml)).
- **WP1: the USM/MPE** continued to maintain the distributed version of the Awe pipelines, running a number of tests using in-house WFI data. In 2006 the Awe system is installed and maintained on 16+1 node Beowulf cluster at MPE, that will act as the main German parallel host in the Awe compute Grid for the reduction of OmegaCam data at the USM/MPE (**WP4**).

- **WP1: Terapix** The performance of the QualityFITS package, developed by Terapix, was compared to the AstroWise system.
- **WP2 Terapix:** The upgrading of the Panorapix viewer was not successful, due to unexpected complications with updating the software libraries used by graphic boards and instead the AstroWise system has been connected to the EuroVO Aladin viewer.
- **WP2 MPE/USM:** The MPE/USM group finished the development and delivered two pieces of software to AstroWise, the photometric redshift code (**photoz**) together with a Python interface, and the Munich Difference Imaging Analysis tool (**MDIA**) which also allows to perform PSF convolution. Corresponding sections to the AstroWise HowTo web pages were also made available.
- **WP2 OAC:** The following new functionalities were added to the **VODIA** software (VST OmegaCAM Difference Image Analysis), and is delivered to the project:
  - output files contain central times and RA/DEC of each exposure
  - The python interface verifies that the sky was not subtracted during SWARP;
  - negative pixels are automatically converted to high positive (saturated) values and are not considered.
  - The user can specify a list of best PSF images to be used to create the reference frame.
  - A new inspect method was created for a quick look of the results including: plot of the light-curve, Discrete Fourier Transform (DFT) and extraction of the main frequencies, computing phases, plot DFT and phase diagram, create small images ("stamps") of each single variable candidate.
- **WP2 NOVA:** The development at NOVA/OmegaCEN of fast computer algorithms to associate (**Associate tools**) on the basis of positional coordinates two very large tables, with at least 100.000 entries (associate tool) was applied in application programmes to compute "global astrometric solutions", i.e. astrometric solutions which make use of information provided by overlapping images. The Global astrometric solutions were further refined and qualified on real data, leading to a low failure rate, less than a few percent, which for most cases were due to the presence of very bright stars.
- **WP2: NOVA/OmegaCEN** maintained and delivered the **dbViewer** (Graphical User Interface) to view the contents of the database via the internet. This web service has been bundled with other Awe web services in the Awe portal, a deliverable of the project allowing users to access the database throughout Europe (and beyond) without requiring software licenses. Publishing data into Euro-VO is prepared.
- **WP2 NOVA: Image cutout server** was built, qualified and delivered. This tool can make customized thumbnails of astronomical sources to be used for inspection or input to source analysis applications
- **WP3:** The federated database for documents and source code, CVS, has been successfully used at the various partner institutes, providing the basic computer network infrastructure. The database populated with source code, documentation and also internet web site data was maintained This system successfully provides the backbone for a common partner wide development network. The web services have been bundled on internet in the Awe-portal pages.

- **WP3 NOVA:** The Oracle database server has been maintained and upgraded at NOVA/OmegaCEN. Also the fileserver was maintained. The layout of the data in the Oracle database was changed to accommodate Advanced Replication which is needed for the federated database. The federated fileserver was upgraded to support authorized deletion of federated files. The federation of database servers at the disposal of the users and connected in real time through Advanced replication was delivered to the project.
- **WP3** Oracle Advanced Replication was used to implement the federated database and a federation between Astro-Wise nodes was established and tested. The federated database is set up at the Astro-Wise nodes on production hardware. The federated fileserver have been connected and are in daily use.
- **WP3 The MPE/USM** node continued to participate in the efforts to implement and test the federation of the AstroWise databases. To achieve this, a test database was federated with Groningen and queries running over tables in both databases and timing tests were carried out. In addition tests of the automatic replication of newly created data were run.
- **WP3 OAC:** Maintenance of Oracle 10g on linux (previous server) and Solaris (new server). Tests on federated DB with Oracle streams were performed.
- **WP3 OAC:** a new DB server was bought and is now running Oracle 10g. The new server is a SUN V40z equipped with two Opteron CPUs, 4 GB of DDR RAM and 500 GB of U320SCSI 10000 rpm disks. Thanks to this new facility the computing time to calculate the global astrometric solution was reduced by a factor >100
- **WP4 - 5: NOVA/OmegaCEN:** Astro-Wise's central CVS repositories, Oracle db and servers with ~20 Tbyte of disk space were maintained. The AstroWise servers are connected via a 24-port GB ethernet switch to the parallel cluster of the University computing centre. The parallel cluster, a 200 node dual core Opteron Linux cluster, was frequently used to bench mark , qualify and operate the AstroWise system. Research oriented data reduction was done on the new cluster using the Awe system and its "distributed processing environment". The "distributed processing environment" was adapted to allow for synchronized jobs on parallel clusters. This is useful for doing global astrometry and processes such as regridding combined with coaddition. The compute grid and storage grid was delivered to the project.
- **WP4 MPE/USM:** After the hardware for (first) remote computing cluster (with 16 double CPUs and 16 TByte of disk) had been delivered by the end of 2005, the cluster was put into operation in the first months of 2006. The AstroWise software with the underlying tools were installed. In May 2006 the definition of a second computing cluster was started with benchmark testings and performance simulations. The system was purchased in November 2006. This second cluster has 124 4-CPU SUN nodes and 100 TByte IBM disk space attached to a PByte robotic storing device. The system will be operated by the Max-Planck Computing Center in Garching and will soon run AstroWise.
- **WP4 and WP5 (OAC):** A new operation system (scientific linux) was installed on the OAC beowulf cluster.
- **WP4 (OAC)** The data storage system was bought and is now available. The system is a NS704 from EMC, equipped with four NAS heads, two storage controllers populated with 12.5 TB (raw) of fibre channel disks. Storage capacity can be scaled up to 130 TB with a very high performance in data delivery (the storage is connected to the Beowulf cluster through 24 Gigabit Ethernet connections divided in four trunked channels).

- **WP5 MPE/USM:** The installation and setup of the first storage system were performed. Benchmark testings and performance simulations for the storing device of the second cluster (see above) were performed to define the system.
- **WP6:** 1 PI meeting was organized, regular quarterly telecons were held and several bi-lateral visits were organized.
- **WP6:** The Astro-Wise consortium continued its collaboration with the “satellite” node Sternwarte University of Bonn. One representative visited the Nov 2006 workshop.
- **WP6:** Meetings with EU-EURO-VO, the EU-EURO-VO-DCA (Data Center Alliance) board and EU-RTN-SISCO were attended by Valentijn for coordination with the virtual observatory projects and the Survey design network.
- **WP6 MPE/USM:** Contact with the others of the AstroWise consortium were held through regular teleconferences, participation in the AstroWise Leiden workshop in November, attended by 6 USM/MPE people, and a visit to Naples.
- **WP6 NOVA:** A one week Astro-Wise Workshop “Preparing for Surveys” was organized at the Lorentz center in Leiden. A two day milestone team meeting was organized at OmegaCEN
- **WP2 Terapix:** The Panorapix visualization tool: F. Magnard, Emmanuel Bertin and Anis Rojbi are now working together on the new version of Panorapix, using OpenGL and Shaders will increase its efficiency and portability. This is the top priority for Terapix for the coming year.
- **WP5 Terapix:** mostly focussed on using the hardware and testing its performances with new Dual-core processors integrated in the heterogeneous Terapix cluster.

## 2.3 Comparison

The contract reads:

The programme is split into three phases, each ending with a deliverable and milestone. Phase I is characterised by design reviews and creating class definitions as soon as possible to receive wide-field imaging data, in order to avoid possible later backwards compatibility problems. During Phase II the system will be developed to prepare for mass production, while during Phase III the system will be fully qualified, populated, and delivered.

Phase I: be ready to receive first data without later re-definitions or re-shuffling

**Table 4.1 Milestones and deliverables - extract for PHASE II - III**

<b>Mile- Stone</b>	<b>Months after kick-off</b>	<b>Description of milestones, actions and deliverables</b>
Phase II: evaluate and prepare for mass production		
9	27	action: upgrade of mass storage WP4- ready for bulk operations
10	30	milestone: system full mass production proof (i.e. 100 Tbyte data volume)  deliverable: operational pipeline, tools, data, results in Tbyte regime
Phase III: qualify, operate and populate the mass production system		
11	33	milestone: review and set goals and schedule for final system
12	48	milestone: end of implementation/development work of the project deliverable: key objective of the project - a qualified survey system to be used for the derivation of survey and research results; including pipeline, software tools and databases.

During the fifth year the programme followed the plan as outlined in Annex II - see table above.

However, already during year three it became clear that ESO will delay its commissioning of the VST telescope and its camera (OmegaCAM).

Since this is one of the most important infrastructures the AstroWise system will support this implies that the foreseen qualification, operation and mass production will have to be delayed to the period December 2005 – November 2006.

Accordingly, the PIs decided to re-schedule the last part of Phase II and the Phase III in a 5 year

Programme that can be completed in November 2006, without requiring additional funding.

An Amendment to the Contract has been set up and was signed by all partners and countersigned by the Commission. In effect the Phase III has started on 18 November 2005, when the portal was launched and the general workshop and tutorial started.

Unfortunately during the fifth year VST was not delivered by ESO, but the system was extensively tested and populated for mass production using OmegaCAM laboratory test data, next to several other data sources such as the [WFI@2.2m](#) and the SuprimeCAM@8m Subaru telescope totalling over 10 TByte of observational data

## 2.4 Planned activities in the next period

EU contract comes to an end, but partners decided to continue support both for operations and maintenance of the system and will bid for funding for support for existing e-science infrastructures for massive data take up and deployment over other communities in the FP7 programme.

### 3 Deliverables

This section should contain the list of contractual deliverables (reports, software codes, experimental results, prototype products, etc.) of any tasks completed to date and a comparison with Annex 1 to the contract.

The following deliverables were maintained and finalised/delivered during the reporting period:

- A fully operational information system for wide field imaging for Europe, ready for lookup or research.
- AstroWise Environment: User and developer manual - update November 2006 (246 pages)
- CVS distributed project documentation
- CVS distributed source code
- Operating image pipeline
- Operating Calibration Pipelines
- Operating web services, including the novel TARGET processor, and the calibration file time stamp manager (CALts).
- Dedicated application packages MDIA, VODIA, GALFIT, GALPHOT and Photo-z.
- Extensive web portal [www.astro-wise.org/portal](http://www.astro-wise.org/portal) including web services and documentation (docs + HowTos).
- Hardware, parallel computers installed at NOVA/OmegaCEN, OAC, USM and Terapix-, together forming a compute GRID
- Hardware, central fileserver and database installed at NOVA/OmegaCEN together forming a data GRID
- A federated and real time synchronized distributed database system
- ASTRO-WISE web site with internal reports on various WP evaluations
- Test benches for database scalability- WP3
- Benchmarks Beowulf clusters WP4
- Hard disk towers for Terabyte storage – WP5
- 30 Tbyte of Wide field imaging data ingested into the AstroWise system .

The software deliverables listed above were delivered to the SternWarte Bonn, who as a satellite node installed the Astro-Wise system.

State whether the research objectives, as set down in Annex I to the contract, are still relevant and achievable. If not explain, why.

Has the methodological approach changed from that described in the contract. If so, how?

Any significant deviations from the planned schedule as described in Annex I to the contract should be clearly highlighted and discussed, stating the recommended actions to correct such deviations.

All objectives have been achieved- the e-science aspects of the project were more strengthened than originally outlined.

The programme has been implemented remarkably close to what has been outlined in the Contract Annex I.

#### 4.0 Exploitation and Dissemination of Results

This section should describe all project aspects that are pertinent to exploitation and dissemination of results, such as:

- publications and conference presentations resulting from the project;
- patentable results, including a list of patents applied for; contacts with potential users and an indication of customer requirements.

List in Annex 1 the most significant publications/conference presentations/patents .

See Annex 1 and <http://www.astro-wise.org> and [www.astro-wise.org/portal](http://www.astro-wise.org/portal)

See list of publications presentations at conferences.

July 2006 Visit of major of Groningen Dr Wallage, rector of University Groningen Prof Zwarts and dean of Faculty Prof Wiersma at Kapteyn institute

Coordinator Valentijn and Dr G. Verdoes Kleijn give a demo of the ASTRO-Wise system





## **5 Management and Coordination**

### **5.0 Meetings**

Meetings attended by AstroWise Staff:

1. Astro-Wise Workshop: "Astro-Wise Preparing for Surveys" at the Lorentz Center in Leiden (see Agenda at [www.astro-wise.org](http://www.astro-wise.org)). 20-23 November 2006 , E.A.Valentijn, K.Kuijken, R.Silvotti (I), , Y. Mellier (F), E. Bertin (F), J.Snigula (D), D. Wilman (D), O-M. Cordes (D), E.Deul, K. Begeman, D.Boxhoorn, W-J Vriend, E.Helmich, R.Vermeij, Ph. Heraudeau, M. Tempelaar, T.Schneider, G.Sikkema, G. Verdoes Kleijn, F.Getman (I), S.Leccia (I), and ,

2. PI meeting 22 Nov 2006 at Leiden minuted at [www.astro-wise.org](http://www.astro-wise.org) , E.A. Valentijn, K. Kuijken, G. Verdoes Kleijn, O.Cordes (Bonn), R. Silvotti (for Cappacioli), D.Wilman (for Bender), Y. Mellier
3. ADASS XVI, 15-18 October 2006, Tucson, Arizona, USA
4. IAU XXXVIth General Assembly, 14-25 August, Praha, Czech Republic
5. AstroWise Milestone Meeting Groningen, NL, 19-20 April 2006
6. visit mayor Groningen Wallage & Dean Rector University Groningen at OmegaCEN – AstroWise demo
7. 2 Nov GigoPart VO presentation Valentijn Utrecht, NL
8. Photometric Standardization workshop, Blankenberge, Belgium, 8-11 May, 2006
9. SISCO network cosmology conference Durham, UK, July 31- August 4, 2006

## 5.1 Coordination activities

The contract reads:

The PM team meets at least two times a year, and is in monthly contact with the local contacts. The project management organisation includes the following ingredients to support the communication between partners spread over different sites:

- weekly teleconferencing on fixed days and times between local contact persons and other team members
- two-monthly meetings reviewing work-packages progress and setting priorities for the next quarter
- 6-monthly project review with PM team and external scientists

There will be an extensive infrastructure channelling, procedurising and defining fixed formats (classes) of all digital information exchanged between partners. Most of this is handled by work-package 3, which also provides an important management tool. To sum up, the exchange of digital information is supported by the following infrastructure:

- project internet web-site, with pages for all official documentation, manuals, arrangements, names of associated persons.  
pass word protected informal pages for intra consortium communications
- federated database for text files, such as source code, manuals etc (eg. CVS)
- federated database for all relevant wide-field imaging data, like calibration files

We have been operating closely along the lines set out in the contract.

## 5.2 Tasks per partner

Detailed up-to-date tables of persons actually working on the various work packages are posted on our webpages. Partners have contributed to the various workpackages in agreement to what was outlined in the contract and the Amendment. The latter involved some shift of work planned at ESO to NOVA.

ASTRO-WISE funded personnel include:

OAC: 1.0 scientific programmers  
 Terapix: \*\* scientific programmer  
 NOVA: 1.8 scientific programmers + 1.0 qualification scientist  
 NOVA-coord: (0.6+0.4) db administrator  
 USM: 1.0 scientific programmer/researcher  
 ESO: no funded personnel

The coordinator maintains detailed sheets (Excel) recording the actual spend human resources for each of the work packages for the various partners, both for EU funded staff and matching staff (eligible costs).

Table 4.2 presents a copy of this Table from the Contract, in which we added in the last three columns, after the /, the totals of the actually spend human resources of each partner during the reporting period.

Due to the three different cost models operated by the various partners any tabulated column has the potential to give a false impression. To avoid misunderstandings we have added in the last column the cost model employed by the particular partner.

The last three columns of Table 4.2 are defined as follows:

*Two but last column:*

Number of man month funded by the project for the whole contract period, as copied from the Contract. After the / we have indicated how much approximately of these **contract units** have been spend by the partners during the contract period.

*One but last column:*

Number of manmonth contributing to the eligible costs (matching costs, having different definitions in the various costmodels) as copied from table 4.2 from the contract, with after the / the total amount of manmonth during the reporting period assigned by partners to the project, but for which no direct EU funding is given, but which are accounted in the Full cost models.

*Last column:*

In the last column we list the actual human resources funded by the project expressed in person hours. Note that in AC models these are the additional costs, while in the full cost models these are the total costs to be multiplied by the fractional EU contribution for deriving the actual EU funding.

The last column is the only column which really matters in the actual financial transactions, while the previous two columns serve to trace the actual costs of the last column to the tables in the contract and partners agreement (Annex II 4.3).

**Table 4.2 Human resources by participant and work package with totals for year 5**

Human resource breakdown (person-month)										
	WP1 Proc	WP2 tools	WP3 db	WP4 Beo	WP5 sto	WP6 cor	Total	Funded by project /year 5 [month]	Not funded by project /year 5 [month]	Funded by project in year 5 [hours]
NOVA	41	74	52	16	9	50	242	66/24	176/21	4355 AC
NOVA/co-or			12			12	24	24/12	0/4	1460 AC
ESO	30	72	18	6	6	8	140	36/0	104/0	0 FC
OAC	33	34	18	7	14	12	118	30/6	88/15	1760 AC
TERAPIX	10	28	27			14	79	48/4	31/7	1510 FF
USM	20	24	3	7	6	8	68	18/7	50/19	1432 AC
Total	134	232	130	36	35	104	671	222/53	449/66	10.617

Note that throughout the year 2006 all the USM/MPE Awe activities were financed with funds outside of the AstroWiseproject, except for the position of Johannes Koppenhoefer. H. Vaith joined the group in July 2007 and dedicated all his time to AstroWise issues.

### 5.3 Budget

An integrated cost statement has been submitted in the Cost Statements in Form E2, which summarizes for each partner the costs for the various budget posts. The consumed money for the fifth year matches well to the budget, which can be seen directly by comparing the global numbers on form E2 with that of form A4 "Yearly costs summary for the first year" on the contract.

Cost-Statement form E3 lists the transferred sums from Coordinator to partners.

## 6. SUPPLEMENTARY INFORMATION

### Annex 1 – List of Publications/Patents

(see <http://www.astro-wise.org/publications.shtml> and <http://www.astro-wise.org/documents.shtml> for full articles)

(1) ASTRO-WISE: " Architectural Design Document" Version 1.0 – Internal Report  
78 pages

(2) ASTRO-WISE:"Users and Programmers Manual, version Nov 2006" 246 pages

(3) Heraudeau, Ph and Valentijn E. A., "An optical survey of the ELAIS-S2 field, data reduction with Astro-WISE" SF2A-2005: Semaine de l'Astrophysique Francaise, meeting

held in Strasbourg, France, June 27 - July 1, 2005, Edited by F. Casoli, T. Contini, J.M. Hameury and L. Pagani. Published by EdP-Sciences, Conference Series, 2005, p. 717

(4) Heraudeau, Ph and Valentijn E. A., 2005, in Proc. The fabulous destiny of galaxies. Bridging past and present, Marseille, 2005, in press.

``An optical survey of the ELAIS-S2 field, data reduction with Astro-WISE'

(5) Alcalá J.M., Marconi M., Ripepi V., de Martino D., Musella I., et al., 2005, "The stellar VST-GTO surveys at the INAF-OA-Capodimonte", in proc. of the XLIX meeting of the Italian Astronomical Society, Mem. Soc. Astron. Ital., in press

(6) Baruffolo A., Alcalá J.M., Benacchio L., et al., 2005, "WP10: Astrophysics in GRID.it", in proc. of "Computational GRIDs for astrophysics in Italy", in press  
(<http://wwwas.oat.ts.astro.it/wiki/bin/view/GridWshop/GridWorkshopPapers>)

(7) Pavlov, M., Alcalá J.M., and Valentijn. E.A., 2005, "VST processing within the GRID" in proc. of "Computational GRIDs for astrophysics in Italy", in press  
(<http://wwwas.oat.ts.astro.it/wiki/bin/view/GridWshop/GridWorkshopPapers>)

(8) Valentijn & Verdoes Kleijn 2006, "The Astro-WISE system: a federated information accumulator for astronomy", ERCIM News, 65, 20

(9) Valentijn et al., 2006, "Astro-WISE: Chaining to the Universe", ADASS XVI, ASP Conference Series, 2006, Shaw, Hill, Bell, eds, in press [2007astro.ph..2189V](#)

(10) Sikkema et al, 2006, "Globular cluster systems of six shell galaxies", Astronomy & Astrophysics, 458, 53

(11) Valentijn & Begeman, 2006, "Creating and analysing multi-dimensional data in Astro-WISE environment: a test case", white paper

(12) Verdoes Kleijn et al., 2006, "The secondary standards programme for OmegaCAM at the VST", Standardization workshop, ASP Conference series, Sterken editor, in press, [2006astro.ph.12469V](#)

(13) Valentijn, Verdoes Kleijn, The Virtual Observatory in Action: New Science, New Technology, and Next Generation Facilities, 26th meeting of the IAU, Special Session 3, 17-18, 21-22 August, 2006 in Prague, Czech Republic, SPS3, #39

(14) Kovac, 2007, "Searching for the lowest mass galaxies: an HI perspective", PhD thesis, Ch 5, University Groningen